Results of the MiniBooNE Oscillation Experiment

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Before MiniBooNE

Before MiniBooNE: The LSND Experiment



Oscillation Status After LSND



This signal looks very different from the others...

- Much higher $\Delta m^2 = 0.1 10 \text{ eV}^2$
- Much smaller mixing angle
- Only one experiment!

Kamioka, IMB, Super K, Soudan II, Macro, K2K $\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$

Homestake, Sage, Gallex, Super-K SNO, KamLAND $\Delta m^2 = 8.2 \times 10^{-5} \text{ eV}^2$

In SM there are only 3 neutrinos



- Three distinct neutrino oscillation signals, with $\Delta m_{solar}^2 + \Delta m_{atm}^2 \neq \Delta m_{LSND}^2$
- For three neutrinos,

expect $\Delta m_{21}^2 + \Delta m_{32}^2 = \Delta m_{31}^2$

How can one get 3 distinct Δm^2 ?

- One of the experimental measurements is wrong
- One of the experimental measurements is not neutrino oscillations
 - Neutrino decay
 - Neutrino production from flavor violating decays
- Additional "sterile" neutrinos involved in oscillations
- CPT violation (or CP viol. and sterile v's) allows different mixing for v's and \overline{v} 's



MiniBooNE

(**Booster Neutrino Experiment**)





Oscillation Analysis

Oscillation Analysis

Extract an oscillation signal.

"Signal range" is approximately 300 MeV < $E_{\rm v}^{\rm QE}$ < 1500 MeV.

One has to: -to minimize background -maximize signal efficiency

Then:

 look for a total excess ("counting exeriment")

 fit for both an excess and energy dependence ("energy fit")



Blind analysis conducted: not looked into signal region("closed box")

Analysis Method

Uses detailed, direct reconstruction of particle tracks, and ratio of fit likelihoods to identify particles.

Apply likelihood fits to three hypotheses:

- -single electron track
- -single muon track

-two electron-like rings (π^0 event hypothesis)

Compare observed light distribution to fit prediction:

Does the track actually look like an electron?

 $log(L_e/L_{\mu})>0$ µ-like events $log(L_e/L_{\mu})>0$ e-like events



Rejecting "µ-like" events

$log(L_e/L_{\mu})>0$ favors electron-like hypothesis



Separation is clean at high energies where muon-like events are long.

This does not separate e/π^0 as photon conversions are electron-like.

Rejecting " π^0 -like" events







Backgrounds

Break-up of the backgrounds

Two main categories of backgrounds: v_{μ} mis-ids and intrinsic v_e



Summary of the backgrounds

Predicted backgrounds passing analysis cuts:

Process	Number of Events
7 CCOF	10
ν_{μ} code	7
Miscellaneous $\nu_{\mu} \in \overline{\nu_{\mu}} \in \mathcal{V}_{\mu}$	13
NC $\pi^{0^{\mu}}$	62
$NC \Delta \rightarrow N\gamma$	20
NC Coherent & Radiative γ	< 1
Dirt Events	17
ν_{e} from μ Decay	132
ν_e from K^+ Decay	71
ν_e from K_L^0 Decay	23
ν_e from π Decay	3
Total Background	358 ± 35(sy st)
$0.26\% \nu_{\mu} \rightarrow \nu_{e}$	163
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The Box Opening: What we found

Open the box and look into $E_v QE$: Return the fit parameters. Is there an oscillation signal?

The Track-based $v_{\mu} \rightarrow v_{e}$ appearance-only result:

Counting Experiment: 475< E, QE <1250 MeV Data: 380 events Expectation: 358 ±19 (stat) ± 35 (sys) events





Best Fit (dashed): (sin²2θ, Δm²) = (0.001, 4 eV²)

Probability of Null Fit: 93% Probability of Best Fit: 99%





Analysis Results

Main Conclusion: The observed reconstructed energy distribution is inconsistent with a $v_{\mu} \rightarrow v_{e}$ 2-neutrino model.



What is New?



Visible energy and angle in E_v QE bins

200<E,<300 MeV

300<E,<475 MeV

475<E_v<3000 MeV



-Low Energy: Excess distributed among visible E, cos ⊕ bins.
-High Energy:Predicted background agrees with data.

Summary of estimated backgrounds vs data

_E, ^{QE} [MeV]	200-300	300-475	475-1250
total background v _e intrinsic v _u induced	284±25 26 258	274±21 67 207	358±35 229 129
[#] ΝC π ⁰	115	76	62
ΝC Δ→Νγ	20	51	20
Dirt	99	50	17
other	24	30	30
Data	375±19	369±19	380±19
Data-MC	91±31	95 ±28	22 <u>+</u> 40

- Low Energy: largest backgrounds a are v_{μ} -induced, in particular: - NC π^{0} - NC $\Delta \rightarrow N\gamma$ - Dirt -High Energy: no significant excess with v_{e} bkgd dominant - Low Energy: largest backgrounds a - NC π^{0} - NC $\Delta \rightarrow N\gamma$ - Dirt - High Energy: no significant excess with v_{e} bkgd dominant - NC π^{0} with photonuclear absorption of π^{0} photon

- new v_{μ} -induced NC photon production (eg: hep-ex:0708.1281v2)

Other data sets available to check signal vs background hypotheses: NuMI neutrinos in MB, MB anti-v run (started JanO6).

Analysis of the events from NuMI beam

4.5 5 E_v[GeV]

NuMI events (for MINOS) detected in MiniBooNE detector!



 ${\bf K}^{\pm} \to \mu^{\pm} \, \nu_{\mu}$

2.5

3

3.5

4

10⁻⁴

10⁻⁵

10⁻⁶

K^{±/0}

0.5

 $\rightarrow \pi \mathbf{e} \mathbf{v}_{\mathbf{e}}$

1.5

2

The beam at MiniBooNE from NuMI is significantly enhanced in v_e from K decay because of the off-axis position.

Analysis of the CCQE events from NuMI beam



Analysis of π^0 events from NuMI beam



Step III: Analysis of v_e events: do we see a similar excess? Search for low energy excess at MiniBooNE with NuMI beam: Ongoing! Please, stay tuned!



MiniBooNE is incompatible with an oscillation $v_{\mu} \rightarrow v_{e}$ appearance-only interpretation of LSND at 98% CL.

Observed deviation of MiniBooNE data from prediction at low energy might be a background: interesting for future $v_u \rightarrow v_e$ searches in same energy region.

It might be a new interesting physics \rightarrow see next talk!

Currently searching for low energy excess at MiniBooNE with NuMI beam with high priority.

MiniBooNE Collaboration

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Testing $e-\pi^0$ separation using the data

